

# Low-energy observables in a SUSY model with spontaneously broken R-Parity

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# Outline

- 1 Introduction
- 2 The considered model with SRPV
- 3 Phenomenological aspects
- 4 Conclusions

# Outline

- 1 Introduction
  - Supersymmetry
  - The MSSM
  - R-Parity
- 2 The considered model with SRPV
- 3 Phenomenological aspects
- 4 Conclusions

## What is Supersymmetry (SUSY)?

- connection between bosons and fermions
- no SUSY particle is observed yet  $\implies$  broken symmetry

## Why do we think about SUSY?

- solution of the hierarchy problem
- dark matter candidates
- unification of gauge couplings at the GUT-scale
- electroweak symmetry breaking (EWSB) automatically
- explanation of neutrino masses

## Minimal Supersymmetric Standard Model (MSSM): particle content

chiral supermultiplets	superfield-notation	spin-0-particles	spin- $\frac{1}{2}$ -particles	quantum number of $SU(3)_C, SU(2)_L, U(1)_Y$
squarks, quarks ( $\times 3$ families)	$\hat{Q}$	$(\tilde{u}_L, \tilde{d}_L)$	$(u_L, d_L)$	$(\mathbf{3}, \mathbf{2}, \frac{1}{6})$
	$\hat{u}^c$	$\tilde{u}_R^*$	$u_R^\dagger$	$(\bar{\mathbf{3}}, \mathbf{1}, -\frac{2}{3})$
	$\hat{d}^c$	$\tilde{d}_R^*$	$d_R^\dagger$	$(\bar{\mathbf{3}}, \mathbf{1}, \frac{1}{3})$
sleptons, leptons ( $\times 3$ families)	$\hat{L}$	$(\tilde{\nu}_e, \tilde{e}_L)$	$(\nu_e, e_L)$	$(\mathbf{1}, \mathbf{2}, -\frac{1}{2})$
	$\hat{e}^c$	$\tilde{e}_R^*$	$e_R^\dagger$	$(\mathbf{1}, \mathbf{1}, 1)$
Higgs, higgsinos	$\hat{H}_u$	$(H_u^+, H_u^0)$	$(\tilde{H}_u^+, \tilde{H}_u^0)$	$(\mathbf{1}, \mathbf{2}, \frac{1}{2})$
	$\hat{H}_d$	$(H_d^0, H_d^-)$	$(\tilde{H}_d^0, \tilde{H}_d^-)$	$(\mathbf{1}, \mathbf{2}, -\frac{1}{2})$

gauge supermultiplets	spin- $\frac{1}{2}$ -particles	spin-1-particles	quantum number of $SU(3)_C, SU(2)_L, U(1)_Y$
gluinos, gluons	$\tilde{g}$	$g$	$(\mathbf{8}, \mathbf{1}, 0)$
winos, $W$ -bosons	$\tilde{W}^\pm, \tilde{W}_3^0$	$W^\pm, W_3^0$	$(\mathbf{1}, \mathbf{3}, 0)$
bino, $B$ -boson	$\tilde{B}$	$B$	$(\mathbf{1}, \mathbf{1}, 0)$

Basic object to build the Lagrangian is the so-called Superpotential:

$$\mathcal{W}_{\text{MSSM}} = \epsilon_{ab} \left[ h_u^{ij} \hat{Q}_i^a \hat{u}_j^c \hat{H}_u^b + h_d^{ij} \hat{Q}_i^b \hat{d}_j^c \hat{H}_d^a + h_e^{ij} \hat{L}_i^b \hat{e}_j^c \hat{H}_d^a - \hat{\mu} \hat{H}_d^a \hat{H}_u^b \right]$$

$\hat{\mu}$  has to be of the order of the electroweak scale  $\mathcal{O}(m_t)$

$\implies \hat{\mu}$ -problem

Sfermion-fermion-gaugino interaction enter the Lagrangian through

$$-\sqrt{2}g (\phi^* T^a \psi) \lambda^a - \sqrt{2}g \lambda^{\dagger a} (\psi^\dagger T^a \phi)$$

Rotating the MSSM-Lagrangian from gauge into mass eigenbasis yields to the following mass spectra of SUSY-particles:

- squarks and sleptons are mixing in pairs
- four charginos  $\chi_1^\pm, \chi_2^\pm$  as result of the mixing between  $\tilde{W}^\pm, \tilde{H}_u^\pm$  and  $\tilde{H}_d^\pm$
- the neutral fermions  $\tilde{W}^0, \tilde{B}, \tilde{H}_u^0$  and  $\tilde{H}_d^0$  form the four neutralinos  $\chi_1^0, \dots, \chi_4^0$

Why do neutrinos not mix with neutralinos in the case of the MSSM?  
 $\implies$  R-Parity is conserved!

## What is R-Parity?

R-Parity is a  $Z_2$ -symmetry.

$$P_R = (-1)^{3(B-L)+2s}$$

$B$  indicates the baryon number,  $L$  the lepton number and  $s$  the spin.

$\implies P = 1$  for SM particles     $\implies P = -1$  for SUSY particles

Conserved R-Parity forbids the terms

$$\mathcal{W}_{\mathbb{R}} = \underbrace{\epsilon_{ab} \left( \frac{1}{2} \lambda_{ijk} \hat{L}_i^a \hat{L}_j^b \hat{e}_k^c + \lambda'_{ijk} \hat{L}_i^a \hat{Q}_j^b \hat{d}_k^c - \varepsilon_i \hat{L}_i^a \hat{H}_u^b \right)}_{\Delta L=1} + \underbrace{\frac{1}{2} \lambda''_{ijk} \hat{u}_i^c \hat{d}_j^c \hat{d}_k^c}_{\Delta B=1}$$



## Why should R-Parity be ...

### ... conserved?

- combined lepton ( $L$ ) and baryon number violation ( $B$ ) give rise to proton decay channels, e.g.  $p \rightarrow \pi^0 e / \pi^+ \nu$
- no contributions to flavour-changing-neutral-currents (FCNC) and lepton-flavour-violation (LFV) on tree-level
- lightest SUSY particale cannot decay  $\implies$  dark matter candidate

### ... broken?

- neutrino-neutralino-mixing  $\implies$  neutrino masses

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- 1 Introduction
- 2 The considered model with SRPV
  - Superpotential
  - Advantages and Effects
  - Introducing neutrino physics
- 3 Phenomenological aspects
- 4 Conclusions

The superpotential reads

$$\mathcal{W} = \epsilon_{ab} \left[ h_u^{ij} \hat{Q}_i^a \hat{u}_j^c \hat{H}_u^b + h_d^{ij} \hat{Q}_i^b \hat{d}_j^c \hat{H}_d^a + h_e^{ij} \hat{L}_i^b \hat{e}_j^c \hat{H}_d^a \right. \\ \left. + h_\nu^{ij} \hat{L}_i^a \hat{\nu}_j^c \hat{H}_u^b - h_0 \hat{H}_d^a \hat{H}_u^b \hat{\Phi} \right] + \frac{\lambda}{3!} \hat{\Phi}^3 + h^{ij} \hat{\Phi} \hat{\nu}_i^c \hat{S}_j$$

A. Masiero and J. W. F. Valle, Phys. Lett. B251, 273-278 (1990)

with additional singlet superfields  $\hat{\Phi}$ ,  $\hat{S}_i$  and  $\hat{\nu}_i^c$  possessing lepton numbers (0, 1, -1). The terms can be identified as:

- known MSSM superpotential terms
- solution of the  $\hat{\mu}$ -problem by an effective  $\mu = h_0 \frac{v_\Phi}{\sqrt{2}}$  since the scalar component of  $\hat{\Phi}$  acquires a VEV (as in NMSSM)
- spontaneously R-Parity violating (SRPV) terms

## Advantages and effects:

- $\mathbb{L}$  as a result of EWSB
- the VEVs  $\langle \tilde{\nu}^c \rangle = \frac{v_R}{\sqrt{2}}$ ,  $\langle \tilde{S} \rangle = \frac{v_S}{\sqrt{2}}$  and  $\langle \tilde{\nu}_{Li} \rangle = \frac{v_{Li}}{\sqrt{2}}$  drive **only**  $\mathbb{L}$   
 $\implies$  no problems with proton decay
- $\hat{\mu}$ -problem is solved by the effective parameter  $\mu = h_0 \frac{v_\Phi}{\sqrt{2}}$
- Due to the broken global symmetry a massless, CP-odd Nambu-Goldstone boson called Majoron arises.
- Neutrino masses are generated at tree level.

In the gauge eigenbasis

$$(\psi^0)^T = ( \tilde{B}^0, \tilde{W}_3^0, \tilde{H}_d^0, \tilde{H}_u^0, \nu^c, S, \Phi, \nu_i )$$

the neutralino mass matrix occurring in  $\mathcal{L}_{\text{neutral}}^{\text{mass}} = -\frac{1}{2} (\psi^0)^T M_N \psi^0 + h.c.$  reads

$$M_N = \begin{pmatrix} M_{\chi^0} & m_{\chi^0 \nu^c} & 0 & m_{\chi^0 \Phi} & m_{\chi^0 \nu} \\ m_{\chi^0 \nu^c}^T & 0 & h \frac{v_\Phi}{\sqrt{2}} & h \frac{v_S}{\sqrt{2}} & m_D^T \\ 0 & h \frac{v_\Phi}{\sqrt{2}} & 0 & h \frac{v_R}{\sqrt{2}} & 0 \\ m_{\chi^0 \Phi}^T & h \frac{v_S}{\sqrt{2}} & h \frac{v_R}{\sqrt{2}} & \lambda \frac{v_\Phi}{\sqrt{2}} & 0 \\ m_{\chi^0 \nu}^T & m_D & 0 & 0 & 0 \end{pmatrix}$$

containing the Dirac mass matrix  $(m_D)_i = \frac{1}{\sqrt{2}} h_\nu^i v_u$  and the R-Parity violating matrix  $(m_{\chi^0 \nu}^T)_i = \left( -\frac{1}{2} g' v_{Li}, \frac{1}{2} g v_{Li}, 0, h_\nu^i \frac{v_R}{\sqrt{2}} \right)_i$

Seesaw-like diagonalization of the neutralino mass matrix leads to the effective neutrino mass matrix

$$m_{\nu\nu}^{\text{eff}} = -M_D^T \cdot M_H^{-1} \cdot M_D,$$

Finally one can find the approximation

$$-\left(m_{\nu\nu}^{\text{eff}}\right)_{ij} = a\Lambda_i\Lambda_j + b(\epsilon_i\Lambda_j + \epsilon_j\Lambda_i) + c\epsilon_i\epsilon_j.$$

including the parameters

$$\Lambda_i = \epsilon_i \mathbf{v}_d + \mu \mathbf{v}_{Li} \quad \text{and} \quad \epsilon_i = \mathbf{h}_\nu^i \frac{\mathbf{v}_R}{\sqrt{2}}.$$

These two vectors are sufficient to control neutrino physics!

Considering **only**  $\vec{\Lambda}$  yields:

$$V^{(\nu)} = \begin{pmatrix} \frac{\sqrt{\Lambda_2^2 + \Lambda_3^2}}{|\vec{\Lambda}|} & 0 & \frac{\Lambda_1}{|\vec{\Lambda}|} \\ -\frac{\Lambda_1 \Lambda_2}{\sqrt{\Lambda_2^2 + \Lambda_3^2} |\vec{\Lambda}|} & \frac{\Lambda_3}{\sqrt{\Lambda_2^2 + \Lambda_3^2}} & \frac{\Lambda_2}{|\vec{\Lambda}|} \\ -\frac{\Lambda_1 \Lambda_3}{\sqrt{\Lambda_2^2 + \Lambda_3^2} |\vec{\Lambda}|} & -\frac{\Lambda_2}{\sqrt{\Lambda_2^2 + \Lambda_3^2}} & \frac{\Lambda_3}{|\vec{\Lambda}|} \end{pmatrix}$$

$\Rightarrow$  two angles for direct reading

$$\tan^2 \theta_R \approx \left( \frac{\Lambda_1}{\sqrt{\Lambda_2^2 + \Lambda_3^2}} \right)^2 \quad \text{and} \quad \tan^2 \theta_{\text{atm}} \approx \left( \frac{\Lambda_2}{\Lambda_3} \right)^2$$

Generation of third angle requires  $\vec{\tilde{\epsilon}} = (V^{(\nu)})^T \cdot \vec{\epsilon}$ :

$$\tan^2 \theta_{\text{sol}} \approx \left( \frac{\tilde{\epsilon}_1}{\tilde{\epsilon}_2} \right)^2$$

Current neutrino oscillation data containing mixing angles and the mass differences  $\Delta m_{ij}^2 = m_i^2 - m_j^2$ :

parameter	best fit ( $1\sigma$ )	$2\sigma$	$3\sigma$
$\Delta m_{21}^2$ $\left[ 10^{-5} \text{ eV}^2 \right]$	$7.65^{+0.23}_{-0.20}$	$7.25 - 8.11$	$7.05 - 8.34$
$\Delta m_{31}^2$ $\left[ 10^{-2} \text{ eV}^2 \right]$	$2.40^{+0.12}_{-0.11}$	$2.18 - 2.64$	$2.07 - 2.75$
$\sin^2 \theta_{12}$	$0.304^{+0.022}_{-0.016}$	$0.27 - 0.35$	$0.25 - 0.37$
$\sin^2 \theta_{23}$	$0.50^{+0.07}_{-0.06}$	$0.39 - 0.63$	$0.36 - 0.67$
$\sin^2 \theta_{13}$	$0.01^{+0.016}_{-0.011}$	$\leq 0.040$	$\leq 0.056$

T. Schwetz, M. Tortola and J. W. F. Valle, arXiv:0808.2016 (2008)

All allowed values can be reached by varying  $\vec{\Lambda}$  and  $\vec{\epsilon}$  respectively the  $\mathbb{R}$  parameters.



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- 3 Phenomenological aspects**
  - Approach and Motivation
  - Inclusive B-Meson decay  $B \rightarrow X_S ll$
  - Leptonic decays
  - Chosen parameter space
  - Numerical results
- 4 Conclusions

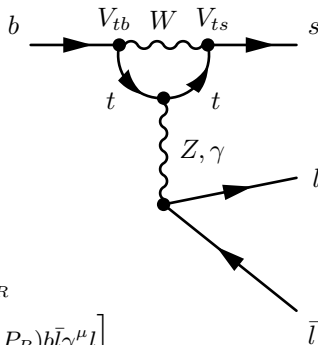
## Approach and Motivation

- fitting the  $\mathcal{R}$  parameters to neutrino data
- searching pocesses which are giving additional constraints (e.g. loop suppressed or forbidden processes)

## $B \rightarrow X_S ll$ in SM

- loop suppressed
- matrix element:

$$\mathcal{M}_{\text{SM}} = \frac{G_F \alpha}{\sqrt{2} \pi} V_{ts}^* V_{tb} \cdot \left[ (C_9^{\text{eff}} - C_{10}) \bar{s}_L \gamma_\mu b_L \bar{l}_L \gamma^\mu l_L + (C_9^{\text{eff}} + C_{10}) \bar{s}_L \gamma_\mu b_L \bar{l}_R \gamma^\mu l_R - 2C_7^{\text{eff}} \bar{s}_i \sigma_{\mu\nu} \frac{q^\nu}{q^2} (m_s P_L + m_b P_R) b \bar{l} \gamma^\mu l \right]$$



- experimental value:

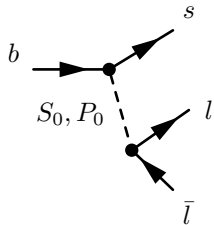
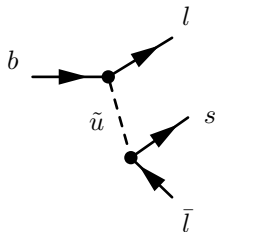
$$\mathcal{B}(B \rightarrow X_S l^+ l^-) = (1.60 \pm 0.51) \cdot 10^{-6}$$

T. Huber et al., Nucl. Phys. B740, 105-137 (2006)

## $B \rightarrow X_S ll$ in SRPV Model

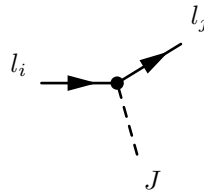
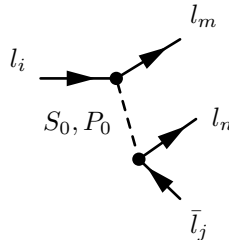
- treelevel process suppressed by  $R$  parameters
- additional contributions:

$$\begin{aligned}
 \mathcal{M}_{\text{add}} = & \frac{g^2}{2m_{\tilde{u}_L}^2} (C_{LR}(\bar{l}_L \gamma^\mu l_L \bar{s}_R \gamma_\mu b_R) \\
 & + C_{RL}(\bar{l}_R \gamma^\mu l_R \bar{s}_L \gamma_\mu b_L) \\
 & + C_{LRLR}(\bar{l}_L l_R \bar{s}_L b_R) \\
 & + C_{RLRL}(\bar{l}_R l_L \bar{s}_R b_L) \\
 & + C_{RLLR}(\bar{l}_R l_L \bar{s}_L b_R) \\
 & + C_{LRRL}(\bar{l}_L l_R \bar{s}_R b_L) \\
 & + \frac{C_T}{8}(\bar{l} \sigma^{\mu\nu} l \bar{s} \sigma_{\mu\nu} b) + i \frac{C_{TE}}{16}(\bar{l} \sigma_{\mu\nu} l \bar{s} \sigma_{\alpha\beta} b) \epsilon^{\mu\nu\alpha\beta} )
 \end{aligned}$$



## Leptonic decays

- violated lepton flavour conservation (LFC)
- not predicted by the SM or MSSM on tree-level



Two mSUGRA scenarios:

SPS1a:

parameter	value in GeV
$m_{1/2}$	250
$m_0$	100
$A_0$	-100

parameter	value
$\text{sign}(\mu)$	+1
$\tan \beta$	10

SPS4:

parameter	value in GeV
$m_{1/2}$	300
$m_0$	400
$A_0$	0

parameter	value
$\text{sign}(\mu)$	+1
$\tan \beta$	50

RGEs of the MSSM are used to get the corresponding MSSM parameters at electroweak scale.

$\mathcal{R}$  violating parameters:

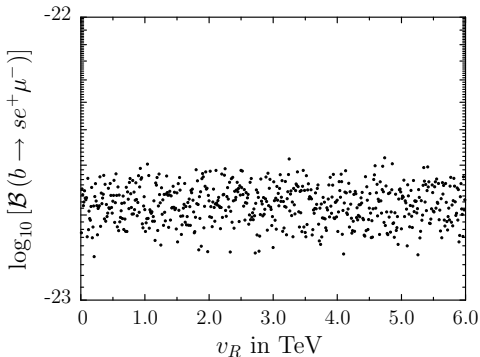
- determined by neutrino fit routine:

$$\Lambda_i = \epsilon_i v_d + \mu v_{Li} \quad \text{and} \quad \epsilon_i = h_\nu^i \frac{v_R}{\sqrt{2}}$$

- fixed by the effective parameter  $\mu$

$$\mu = h_0 \frac{v_\Phi}{\sqrt{2}}$$

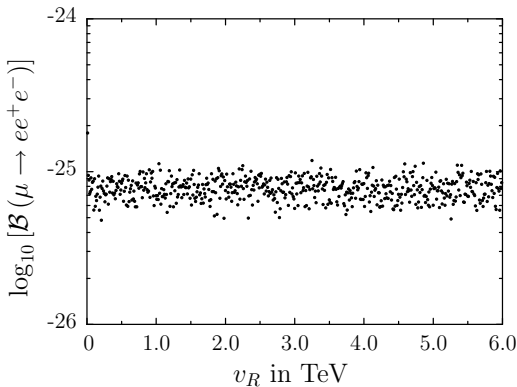
- all variations show only very poor dependences
- $\mathcal{B}(b \rightarrow sl_i^+ l_i^-)$ : almost constant at  $1.92 \cdot 10^{-6}$
- $\mathcal{B}(b \rightarrow sl_i^+ l_j^-)$ : unmeasurable small


 SPS1a:  $\vec{\Lambda} \rightarrow \theta_{\text{atm}}$ , variation of  $v_R = v_S$ 

$\Rightarrow$  no additional constraints to  $\mathcal{R}$  parameters



## Leptonic threebody decays



SPS4:  $\vec{\Lambda} \rightarrow \theta_{\text{atm}}$ , variation of  $v_R = v_S$

All variations show only very poor dependencies on  $\mathcal{R}$  parameters and are much below experimental bounds if measurable at all.

maximal magnitude of the branching ratios:

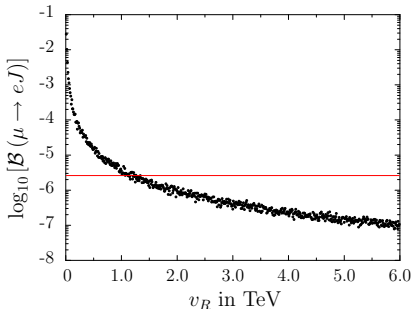
branching ratio	SPS1a		SPS4		experimental upper bound
	$\theta_{\text{atm}} \rightarrow \vec{\Lambda}$	$\theta_{\text{atm}} \rightarrow \vec{\epsilon}$	$\theta_{\text{atm}} \rightarrow \vec{\Lambda}$	$\theta_{\text{atm}} \rightarrow \vec{\epsilon}$	
$\mathcal{B}(\mu \rightarrow ee^+e^-)$	$\mathcal{O}(10^{-26})$	$\mathcal{O}(10^{-27})$	$\mathcal{O}(10^{-25})$	$\mathcal{O}(10^{-23})$	$1.0 \cdot 10^{-12}$
$\mathcal{B}(\tau^- \rightarrow e^-e^+e^-)$	$\mathcal{O}(10^{-24})$	$\mathcal{O}(10^{-25})$	$\mathcal{O}(10^{-24})$	$\mathcal{O}(10^{-21})$	$2.0 \cdot 10^{-7}$
$\mathcal{B}(\tau^- \rightarrow e^- \mu^+ e^-)$	$\mathcal{O}(10^{-24})$	$\mathcal{O}(10^{-32})$	$\mathcal{O}(10^{-24})$	$\mathcal{O}(10^{-21})$	$1.1 \cdot 10^{-7}$
$\mathcal{B}(\tau^- \rightarrow e^- \mu^- e^+)$	$\mathcal{O}(10^{-23})$	$\mathcal{O}(10^{-24})$	$\mathcal{O}(10^{-23})$	$\mathcal{O}(10^{-20})$	$2.7 \cdot 10^{-7}$
$\mathcal{B}(\tau^- \rightarrow e^- \mu^+ \mu^-)$	$\mathcal{O}(10^{-20})$	$\mathcal{O}(10^{-21})$	$\mathcal{O}(10^{-19})$	$\mathcal{O}(10^{-16})$	$3.3 \cdot 10^{-7}$
$\mathcal{B}(\tau^- \rightarrow \mu^- e^+ \mu^-)$	$\mathcal{O}(10^{-23})$	$\mathcal{O}(10^{-24})$	$\mathcal{O}(10^{-23})$	$\mathcal{O}(10^{-20})$	$1.3 \cdot 10^{-7}$
$\mathcal{B}(\tau^- \rightarrow \mu^- \mu^+ \mu^-)$	$\mathcal{O}(10^{-19})$	$\mathcal{O}(10^{-20})$	$\mathcal{O}(10^{-19})$	$\mathcal{O}(10^{-16})$	$1.9 \cdot 10^{-7}$

U. Bellgardt et al., Nucl. Phys. B299, 1 (1988)

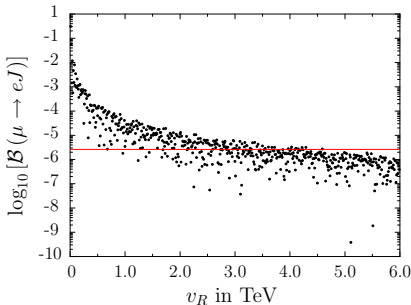
B. Aubert et al., Phys. Rev. Lett. 92, 121801 (2004)

$\Rightarrow$  no additional constraints to  $\mathcal{R}$  parameters

## Leptonic decay $\mu \rightarrow eJ$



SPS1:  $\vec{\Lambda} \rightarrow \theta_{\text{atm}}$ , variation of  $v_R = v_S$



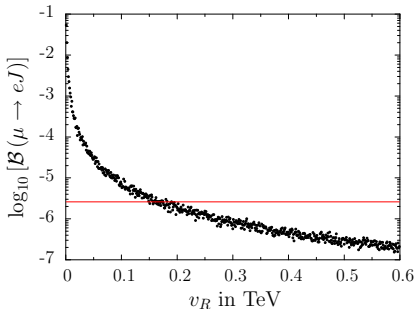
SPS1:  $\vec{\epsilon} \rightarrow \theta_{\text{atm}}$ , variation of  $v_R = v_S$

experimental bound  $2.6 \cdot 10^{-6}$  (TRIUMF)  
 $\Rightarrow v_R$  must be at least about 1 TeV

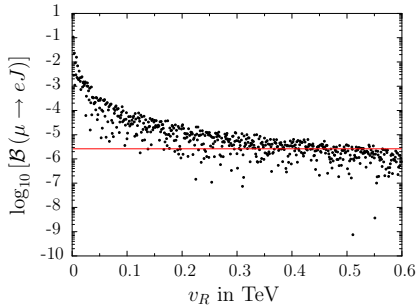
J. C. Romão et al., Nucl. Phys. B363, 369-384 (1991)

A. Jodidio et al., Phys. Rev. D34, 1967-1990 (1986)

## Leptonic decay $\mu \rightarrow eJ$



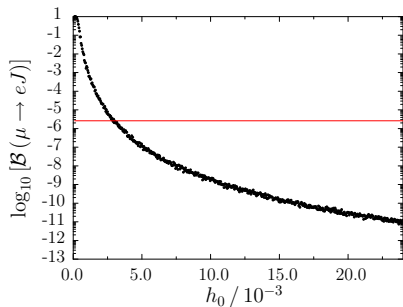
SPS1:  $\vec{\Lambda} \rightarrow \theta_{\text{atm}}$ , variation of  $v_R = 0.1 \cdot v_S$



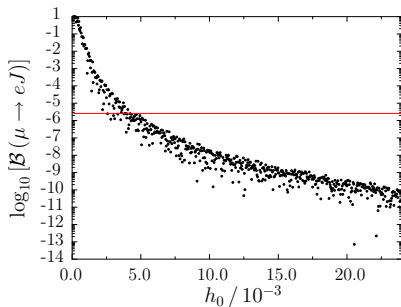
SPS1:  $\vec{e} \rightarrow \theta_{\text{atm}}$ , variation of  $v_R = 0.1 \cdot v_S$

experimental bound  $2.6 \cdot 10^{-6}$  (TRIUMF)  
 $\Rightarrow v_R$  must be at least about 100 GeV

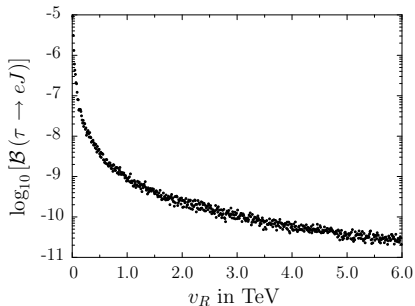
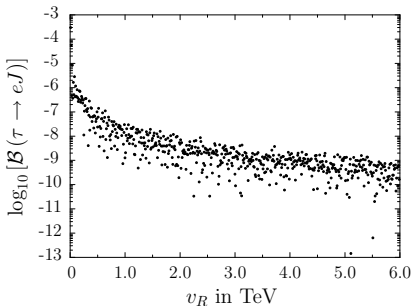
## Leptonic decay $\mu \rightarrow eJ$



SPS1:  $\tilde{L} \rightarrow \theta_{\text{atm}}$ , variation of  $h_0$  by  $v_R = 1$  TeV



SPS1:  $\tilde{e} \rightarrow \theta_{\text{atm}}$ , variation of  $h_0$  by  $v_R = 1$  TeV

Leptonic decay  $\tau \rightarrow eJ$ SPS1:  $\vec{\Lambda} \rightarrow \theta_{\text{atm}}$ , variation of  $v_R = v_S$ SPS1:  $\vec{\epsilon} \rightarrow \theta_{\text{atm}}$ , variation of  $v_R = v_S$ 

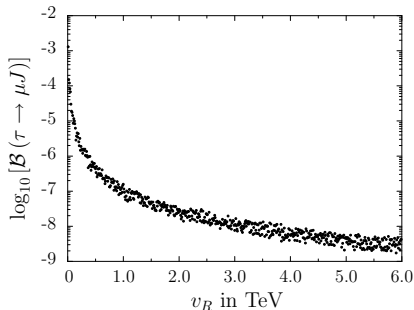
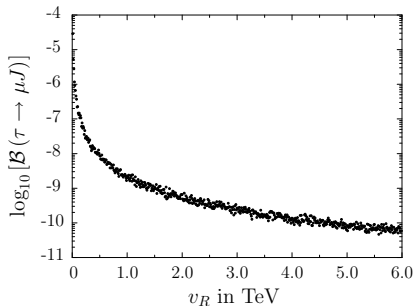
experimental bound  $7.12 \cdot 10^{-3}$  (MARK-III)

$\Rightarrow$  no additional constraints

J. C. Romão et al., Nucl. Phys. B363, 369-384 (1991)

R. M. Baltrusaitis et al., Phys. Rev. Lett. 55, 1842-1845 (1985)

## Leptonic decay $\tau \rightarrow \mu J$



experimental bound  $2.25 \cdot 10^{-2}$  (MARK-III)  
 $\Rightarrow$  no additional constraints

J. C. Romão et al., Nucl. Phys. B363, 369-384 (1991)

R. M. Baltrusaitis et al., Phys. Rev. Lett. 55, 1842-1845 (1985)

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## Conclusions

- $B \rightarrow X_s ll$  and leptonic threebody decays obtain no additional constraints to  $\mathcal{R}$  parameters once neutrino physics are fulfilled
- $\mu \rightarrow eJ$  constricts the parameter space
- $\tau \rightarrow lJ$  yield no crucial bounds yet